

# PX156 Quantum Phenomena and Particle Physics

## Part A - Quantum

### I- Blackbody Radiation

- > Blackbody - an ideal absorber that absorbs all radiation that strikes it
  - also an ideal emitter
- > Blackbody radiation - Continuous Spectrum radiation emitted by an ideal blackbody

→ Stefan-Boltzmann law

$\sigma = \text{Stefan-Boltzmann constant}$

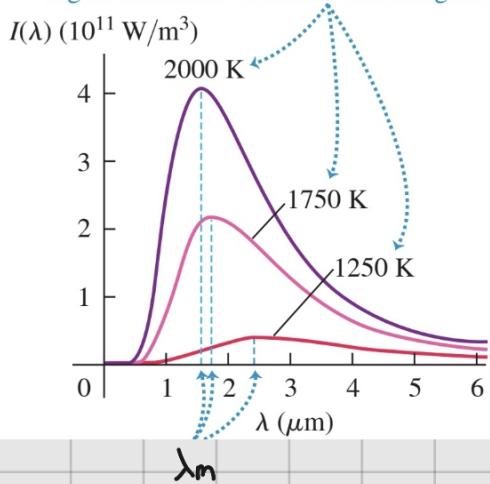
$$I = \sigma T^4$$

- ↳ This is not uniformly distributed over all wavelengths
  - Useful to use Spectral emittance -  $I(\lambda)$  where total intensity

$$I = \int_0^\infty I(\lambda) d\lambda$$

- > Spectral emittance - power per unit wavelength interval
    - Units  $\text{W m}^{-3}$
- ↳ **not the same thing as Intensity**

As the temperature increases, the peak of the spectral emittance curve becomes higher and shifts to shorter wavelengths.



Wien's displacement law:

$$\lambda_{\max} T = 2.898 \times 10^{-3} \text{ mK}$$

$$I(\lambda) = \frac{2\pi c k_B T}{\lambda^4}$$

- The Rayleigh-Jeans law was not derived experimentally

- Agrees with experimental results at high wavelengths
- As  $\lambda \rightarrow 0$ ,  $I(\lambda) \rightarrow \infty$ , not 0 as experimentally found  
↳ the ultraviolet catastrophe

Planck derived a model for BB radiation:

- Energy quantised to  $E_n = nhf \quad n \in \mathbb{Z}$

$$\langle E \rangle = \frac{hf}{\exp(hf/k_B T) - 1}$$

$$I(\lambda) = \frac{2\pi c^2 h}{\lambda^5} \frac{1}{\exp(hc/\lambda k_B T) - 1}$$

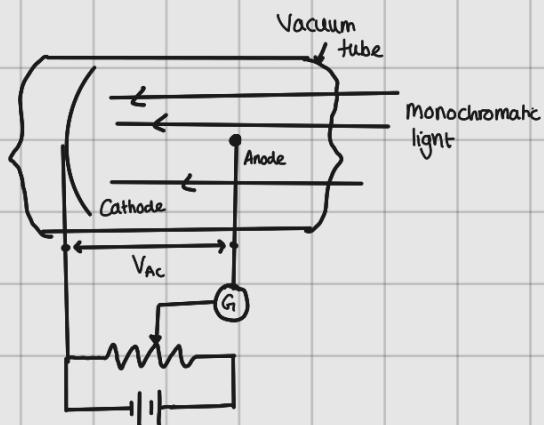
→ Planck's radiation law

## 2 - The Photoelectric Effect

$$E = hf \rightarrow \text{photon energy}$$

$$E_{k\max} = hf - \varphi \rightarrow \text{Photoelectron maximum kinetic energy}$$

> Work function,  $\varphi$  - minimum energy required to remove electrons from a metal's surface



### Experimental observations

- Electrons emitted with range of KE
- Maximum KE dependent on light frequency and independent of light intensity
- No electrons emitted if frequency is below a threshold value

Considering a photon as a relativistic particle:

$$\begin{aligned} E_T &= KE + m_0 c^2 \\ &= (\gamma - 1) m_0 c^2 + m_0 c^2 = \gamma m_0 c^2 = \frac{m_0 c^2}{\sqrt{1 - \frac{u^2}{c^2}}} \end{aligned}$$

as  $u \rightarrow c$ ,  $E \rightarrow \infty$  if  $m_0 \neq 0$ , but  $E = hf$  so  $m_0 = 0$

AS  $m_0 = 0$ ,  $E = pc \rightarrow$

$$p = \frac{E}{c} = \frac{hf}{c} = \frac{h}{\lambda}$$

### 3- Compton Scattering

- A beam of x-rays is fired at a solid target and the wavelength of scattered radiation is measured

$$\lambda' - \lambda = \frac{h}{m_ec} (1 - \cos \theta)$$

$\lambda'$  - Wavelength of Scattered radiation  
 $\lambda$  - Wavelength of incident radiation  
 $\theta$  = Scattering angle

Example:



A photon scatters off a stationary electron. If the photon emerges at an angle of  $60^\circ$  to its initial direction and its wavelength of scattering is doubled, calculate the angle at which the electron recoils,  $\phi$ :

$$\theta = 60^\circ \quad \lambda' = 2\lambda \quad p' = \frac{p}{2} \quad \leftarrow \text{from } p = \frac{h}{\lambda}$$

$$\lambda' - \lambda = \frac{h}{m_ec} (1 - \cos \theta)$$

$$2\lambda - \lambda = \lambda_c (1 - \cos 60^\circ)$$

$$\lambda = \frac{h}{2mc} = \frac{1}{2} \lambda_c$$

$$\text{horizontal: } p = p' \cos\theta + p_e \cos\phi = \frac{p}{2} \cos\theta + p_e \cos\phi$$

$$\text{Vertical: } p' \sin\theta = \frac{p}{2} \sin\theta = p_e \sin\phi$$

$$\frac{p_e \cos\phi}{p_e \sin\phi} = \cot\phi = \frac{p(1 - \frac{1}{2}\cos\theta)}{p/2 \sin\theta} = \frac{2 - \cos\theta}{\sin\theta}$$

$$\text{For } \theta = 60^\circ : \tan\phi = \frac{\sin 60^\circ}{2 - \cos 60^\circ} = \frac{1}{\sqrt{3}} \quad \phi = 30^\circ$$

## 4 - Atomic Spectra

Bohr's model of the atom:

- Electrons orbit the nucleus at discrete distances  $r_n$

$$r_n = \frac{n^2 \epsilon_0 h^2}{\pi m_e e^2}$$

- Electrons lose/gain energy by emitting/absorbing photons, which allows them to move between energy levels

The angular momentum of electrons,  $L$ , is limited by

$$L = n\hbar$$

The total energy for the  $n$ th orbit of hydrogen is

$$E_n = -\frac{hcR}{n^2} = -\frac{13.6 \text{ eV}}{n^2}$$

The wavelength of a photon emitted in a transition from state  $n_1$  to  $n_2$

$$\frac{1}{\lambda} = R \left( \frac{1}{n_2^2} - \frac{1}{n_1^2} \right)$$

## 5- de Broglie Wavelength

- The deBroglie wavelength of a particle is

$$\lambda = \frac{h}{mv} = \frac{h}{p}$$

- Bragg Scattering is the diffraction of x-rays in atomic spacing

$$n\lambda = 2d \sin\theta$$

## 6- Schrödinger's Equation

For a free particle of mass  $m$ :

$$-\frac{\hbar^2}{2m} \frac{\partial^2 \psi}{\partial x^2} = i\hbar \frac{\partial \psi}{\partial t}$$

where  $\psi$  is the wave function, representing the probability of finding a particle at a point in Space at a certain time

↳ A plane wave  $\psi = e^{ikx - i\omega t}$

Satisfies this

Darcie Hurley 2023

The relationship between  $\omega$  and  $k$  can be derived to be

$$\hbar\omega = \frac{\hbar^2 k^2}{2m}$$

The probability density function can be introduced:

$$|\psi(x,t)|^2 = \psi(x,t) \psi^*(x,t)$$

The normalisation function to give the probability of finding the particle in all given locations in the considered space

$$\int |\psi(x,t)|^2 dx = 1$$

> Wave packet - Superposition of many plane waves with an average wavenumber of  $k_0$

L by superposing many waves, one wave is constructed with one maximum of  $|\psi(x,t)|^2$

- This looks like a particle in that it is localised in space

A wave packet can be represented as

$$\psi = \int a(k) e^{ikx - i\omega(k)t} dk$$

where  $a(k)$  is  $k$ -dependent amplitude



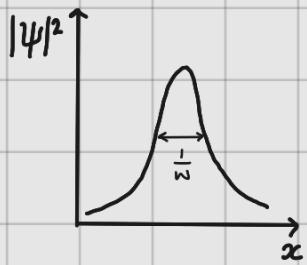
$w$  is the full width at half maximum

$$w = 2\sqrt{\ln 2} \sigma$$

Where the distribution is narrow around  $k_0$ ,  $\psi$  is given by

$$\psi = \exp(ik_0x - i\omega_0 t) \underbrace{\int a(k) \exp(i\Delta k(x - \omega_0't)) dk}_{f(s) = f(x - \omega_0't)}$$

Which represents a travelling wave with  $Vg = \omega_0$



For a wave with width  $w$  in  $k$ , a top-hat distribution can be used

$$a(k) \begin{cases} = A & k_0 - \frac{w}{2} < k < k_0 + \frac{w}{2} \\ = 0 & \text{otherwise} \end{cases}$$

### Time Independent Schrödinger Equation

If a particle is not free, it has a potential energy  $V(x)$ , and Schrödinger's equation becomes

$$i\hbar \frac{\partial}{\partial t} \psi(x, t) = -\frac{\hbar^2}{2m} \frac{\partial^2}{\partial x^2} \psi(x, t) + V(x) \psi(x, t)$$

by letting  $\psi(x, t) = \phi(x) \rho(t)$

$$\rightarrow i\hbar \frac{\partial}{\partial t} (\phi(x) \rho(t)) = -\frac{\hbar^2}{2m} \frac{\partial^2}{\partial x^2} (\phi(x) \rho(t)) + V(x) \phi(x) \rho(t)$$

$$i\hbar \phi(x) \frac{d}{dt} \rho(t) = -\frac{\hbar^2}{2m} \rho(t) \frac{d^2 \phi(x)}{dx^2} + V(x) \phi(x) \rho(t)$$

Dividing by  $\phi(x) \rho(t)$

$$\frac{i\hbar}{\rho} \frac{d\rho}{dt} = -\frac{\hbar^2}{2m\phi} \frac{d^2\phi}{dx^2} + V(x) = E \quad \text{where } E \text{ is constant}$$

$i\hbar \frac{dp}{dt} = Ep$  has solution  $\rho(t) = \exp(-iEt/\hbar)$

$$\boxed{-\frac{\hbar^2}{2m} \frac{d^2\phi}{dx^2} + V\phi = E\phi}$$

→ Solution depends on precise form of  $V(x)$

The probability density of the TISE is given by:

$$|\psi|^2 = \phi \phi^* \rho \rho^*$$

- this is independent of time as  $\rho$  is essentially a phase term
- This is a Stationary State

## 7 - Heisenberg's Uncertainty Principle

$$\Delta x \Delta p \geq \frac{\hbar}{2}$$

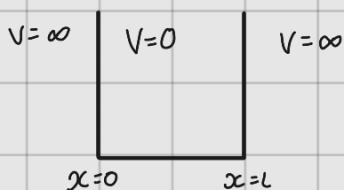
$$\Delta x = \langle x^2 \rangle - \langle x \rangle^2$$

It is impossible to simultaneously determine both the position and momentum of a particle with arbitrarily great precision

## 8 - Potential Wells

> Potential Well - The region surrounding a local minimum of potential energy

### Infinite Potential Well



$$V(x) = \infty \text{ for } x < 0 \text{ and } x > L$$
$$V(x) = 0 \text{ for } 0 \leq x \leq L$$

So that the particle cannot exist outside the well with infinite energy  
 $\phi(x) = 0$  for  $x < 0$  and  $x > L$

For  $0 \leq x \leq L$   $V(x) = 0$  So  $\frac{-\hbar^2}{2m} \frac{d^2\phi}{dx^2} = E\phi$

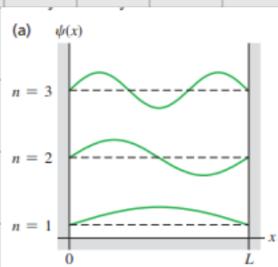
↳  $\frac{d^2\phi}{dx^2} + \frac{2mE}{\hbar^2}\phi = 0$

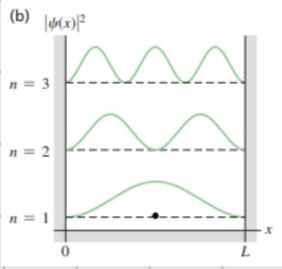
↳  $\frac{d^2\phi}{dx^2} + k^2\phi = 0$  where  $k^2 = \frac{2mE}{\hbar^2}$

Which has solution  $\phi = A \sin(kx) + B \cos(kx)$

- Using boundary conditions  $\phi(0) = 0$  and  $\phi(L) = 0$ ,  $B = 0$ ,  $k = \frac{n\pi}{L}$   
So  $\phi(x) = A \sin(n\pi x/L)$

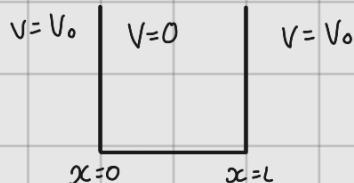
- Use normalisation for  $\int_0^L (A \sin(n\pi x/L))^2 dx = 1$  to find  $A = \sqrt{\frac{2}{L}}$





The value of  $\int |\psi(x)|^2 dx$  at each point represents the probability of finding the particle in small interval  $dx$  about the point

## Finite Potential Well



$$V(x) = V_0 \text{ for } x < 0 \text{ and } x > L$$

$$V(x) = 0 \text{ for } 0 \leq x \leq L$$

$$\text{TISE: } -\frac{\hbar^2}{2m} \frac{d^2\phi}{dx^2} + V_0 \phi = E \phi$$

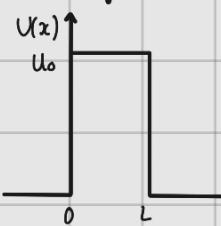
- Inside the well  $\frac{d^2\phi}{dx^2} + k^2\phi = 0$   $k^2 = \frac{2mE}{\hbar^2}$

- Outside the well  $\frac{d^2\phi}{dx^2} - \frac{2m}{\hbar^2} (V_0 - E)\phi = 0$  ← Cannot be solved analytically

## 9 - Potential Barrier

> Potential barrier - A potential energy function with a local maximum

- In classical mechanics, if a particle has total mechanical energy less than the maximum energy of the barrier, it cannot pass through
- In quantum mechanics, the particle may pass through → tunnelling



→ The functions in and out of the barrier must join smoothly.

So  $\psi(x)$  and  $\frac{d\psi}{dx}$  must be continuous at  $x=0, L$

- Before the barrier  $\frac{d^2\phi_1}{dx^2} + \frac{2mE}{\hbar^2}\phi_1 = 0$  with  $\phi_1 = Ae^{ikx} + Be^{-ikx}$
- After the barrier  $\frac{d^2\phi_3}{dx^2} + \frac{2mE}{\hbar^2}\phi_3 = 0$  with  $\phi_3 = Fe^{ikx} + Ge^{-ikx}$
- Inside the barrier  $\frac{d^2\phi_2}{dx^2} + \frac{2m(\varepsilon - U_0)}{\hbar^2}\phi_2 = 0$  with  $\phi_2 = Ce^{ik'x} + De^{-ik'x}$   
where  $k^2 = \frac{2m\varepsilon}{\hbar^2}$  and  $k' = \sqrt{\frac{2m(\varepsilon - U_0)}{\hbar^2}}$

At boundary conditions:

$$x=0: \quad \phi_1 = \phi_2 \quad \frac{d\phi_1}{dx} = \frac{d\phi_2}{dx}$$

$$x=L: \quad \phi_3 = \phi_2 \quad \frac{d\phi_3}{dx} = \frac{d\phi_2}{dx}$$

The transmission probability is given by

$$T \propto e^{2k'L}$$

## 10 - Alpha Decay

$$\frac{dN}{dt} = -\lambda N$$

$$N(t) = N(0) e^{-\lambda t}$$

N - number of nuclei remaining

$\lambda$  - decay rate - probability per second

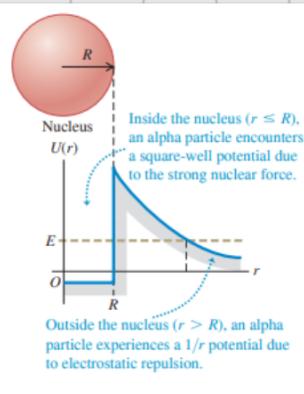
$$\tau = \frac{1}{\lambda}$$

$$t_{1/2} = \frac{\ln 2}{\lambda}$$

$\rightarrow t_{1/2}$  = half life

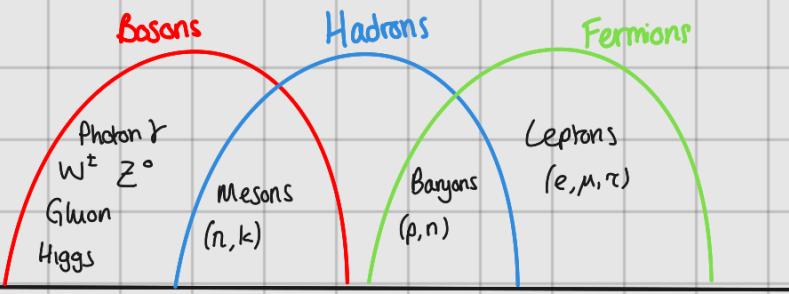
$\rightarrow \tau$  = particle lifetime

In  $\alpha$  decay, the  $\alpha$  particle tunnels through a potential barrier



$\rightarrow$  If an  $\alpha$  inside the nucleus has  $E > 0$ , it can escape the nucleus

## PARTICLES:



### Bosons

- Force associated/carriers
- $n\hbar$  Spin  $n \in \mathbb{Z}$   
( $n=0$  for Higgs)

### Hadrons

- Particles that experience SNF (Comprised of quarks)
- 0 net colour charge and integer electric charge

### Fermions

- Matter particles
- Spin  $\frac{n\hbar}{2}$  for odd  $n$
- 2 classes, each with 3 generations, quarks and leptons

Quantum Numbers: Baryon, Le  $L_\mu L_\tau$ , electric charge, weak isospin, flavour #, energy\*, momentum\*, Colour, quark flavour

= Conserved by all forces

Quark	Flavour #
u	+1/2
d	-1/2
c, t	+1
s, b	-1

Particle	$T_W$
u, c, t	+1/2
d, s, b	-1/2
e, μ, τ	-1/2
$\nu_e, \nu_\mu, \nu_\tau$	+1/2

\*  $m_f \leq m_i$

Weak Isospin:

- $W^\pm$  carries  $T_W = \pm 1$ , so interactions

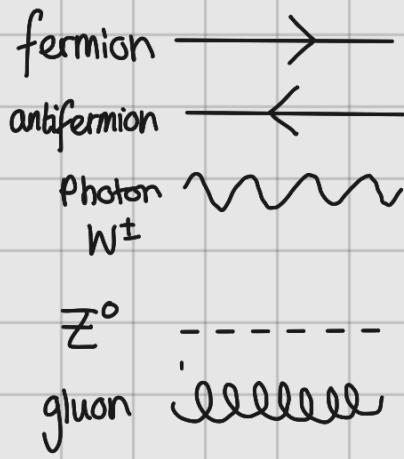
change  $T_W$  of lepton by 1

↑      ↑  
opposite for (right-handed)  
antiparticles

### Colour:

- Quarks have colour charge
- r, g, b  $\rightarrow \bar{r}\bar{r} = \text{Colourless}$        $r\bar{g}b = \text{Colourless}$

## FEYNMANN DIAGRAMS



- gluons only emitted by quarks + gluons
- photons emitted/absorbed by leptons + quarks
- $Z^0$  emitted/absorbed by everything except gluons

- $W$  transforms within lepton generations
- $W$  transforms cross-generationally for quarks, but only changing charge
- No flavour changing neutral interactions

Hierarchy of force choice: Strong → EM → Weak (fastest → slowest)

Crossing Symmetry: Particles can move from  $i \leftrightarrow f$  by turning into anti-particles

$$\hookleftarrow u \rightarrow W^+ + d \quad \text{or} \quad u + W^- \rightarrow d \quad \text{or} \quad W^- \rightarrow d + \bar{u} \dots$$

Changes:

- $Z^0/\gamma/g \leftrightarrow q, \bar{q}$
- $Z^0 \leftrightarrow \nu, \bar{\nu}$
- $Z^0/\gamma \leftrightarrow l^+, l^-$
- $W^\pm \leftrightarrow q, \bar{q}_2$
- $W^\pm \leftrightarrow l \nu$
- Net flavour change  $\rightarrow W^\pm$

Showing Virtual particles have no mass:



- $E_i$  and  $p_i$  fixed, So 4 degrees of freedom
- $E_i = (\rho_i c)^2 + (m_e c^2)^2$  So 3 dof
- $E_i = E_\gamma + E_f$  So 2 dof
- $p_i = p_f + p_\gamma$  So 1 dof
- $E_f = (\rho_f c)^2 + (m_e c^2)^2$  So 0 dof
- $E_\gamma = \rho_\gamma c \rightarrow m_\gamma = 0$

Trying 2 Separate 2 quarks

- Separation  $\uparrow$  = energy density in gluon field  $\uparrow$

- At some point  $E$  is enough to create 2 new quarks

## FORCES:

### Strong

- Acts on particles with colour charge
- Exchange particle is gluon
  - $m, q, T_w = 0$ , Colour =  $C_1, \bar{C}_2$
- Weak at short distances, strong at large
  - Short  $r$ ,  $F \sim 1/r^2$  like  $E_M$
  - long  $r$ ,  $F \sim -k$  ( $k = 1 \text{ GeV/fm}$ )
- Fastest  $\sim 10^{-20} \text{ s} \gg$
- Only force that conserves colour
- conserves quark flavor

### Weak

- Acts on particles with weak isospin
- Exchange particle is  $W^\pm$  for charged,  $Z^0$  for neutral
  - $W$  can couple to  $W$  or  $Z^0$
- Responsible for decay
- Short range:  $F = \alpha_W \frac{e^{-r/R}}{r^2}$
- Slowest  $\sim 10^2 - 10^9 \text{ s}$
- doesn't conserve quark flavor

### Electromagnetism

- Exchange particle is photon,  $\gamma$ 
  - No  $m, q$ ; doesn't interact with  $\gamma$
- Acts between electrically charged particles
- All quantum numbers conserved
- Infinite range  $F \sim 1/r^2$
- Second fastest  $\sim 10^{-15} \text{ s}$
- conserves all quantum numbers

### Charged Current

- Carried by  $W^\pm$  boson
  - $W^\pm: T_w, q_r = \pm 1$
  - heavy
  - only boson to change flavour of  $q + \ell$

### Neutral Current

- Carried by  $Z^0$ 
  - $Z^0: T_w, q_r = 0$
  - heavier than  $W^\pm$
- Only couples particles with their  $\bar{\rho}$   
LO NO FCNC

### Gravity

- Exchange particle is Higgs boson

# SYMMETRY

## Continuous:

- > Symmetries involving a continuous change in some parameter defining the geometry of the system

Translation in

Conserved quantity:

\* See Noether's theorem

- Space  $\vec{x} \rightarrow \vec{x} + \vec{k}$   $\rightarrow$  linear momentum
- angle  $\theta \rightarrow \theta + \theta_0$   $\rightarrow$  angular momentum
- time  $t \rightarrow t + t_0$   $\rightarrow$  energy

↳ laws of physics symmetric under space-time symmetry

## Internal:

- > An operation on the mathematics describing a system

Gauge Symmetry

- Wavefunction phase  $\rightarrow$  electric charge
- Upness & Downness  $\rightarrow$  electroweak charge
- Colour  $\rightarrow$  colour

Conservation of:



$$|\psi|^2 = \psi\psi^*$$

$$\text{If } \psi' = e^{ix}\psi:$$

$$|\psi'|^2 = \psi'\psi'^* = (e^{ix}\psi)(e^{-ix}\psi^*) \\ = \psi\psi^* = |\psi|^2$$

## Discrete:

- > Symmetries based on non-continuous changes to the system

- Parity,  $\hat{P}$  - Spatial inversion ( $x \rightarrow -x$ )  $\leftarrow$   $L = \vec{r} \times \vec{p}$  but after  $\hat{P}$ ,  $\vec{r} = -\vec{r}$   $\vec{p} = -\vec{p}$   
So  $L' = -\vec{r} \times -\vec{p} = L$
- Charge Conjugation,  $\hat{C}$  - Quantum sign inversion ( $p \rightarrow \bar{p}$ )
- Time reversal,  $\hat{T}$  - time inversion ( $t \rightarrow -t$ )
- CP ( $x \rightarrow -x$ ,  $p \rightarrow \bar{p}$ )
- CPT ( $x \rightarrow -x$ ,  $p \rightarrow \bar{p}$ ,  $t \rightarrow -t$ )

↳ All are conserved by Strong & EM, but only CPT by weak too

## DECAYS

Fermi's Golden Rule:  $P(i \rightarrow f) \propto (\text{Amplitude}) \times (\text{Phase Space})$

> Phase Space - tells you how many ways you can distribute energy from the initial state into KE once the final state masses are taken into account

$$N(t) = N(0) e^{-t/\tau}$$

$$\frac{dN(t)}{dt} = -N(t) \Gamma$$

$\tau$  = Particle lifetime  
 $\Gamma = \frac{1}{\tau}$  = decay rate

- Particles have a number of modes to decay. Each mode  $i$  has its own decay rate  $\Gamma_i$ :

$$\Gamma_T = \sum \Gamma_i$$

$$BR_i = \frac{\Gamma_i}{\Gamma_T}$$

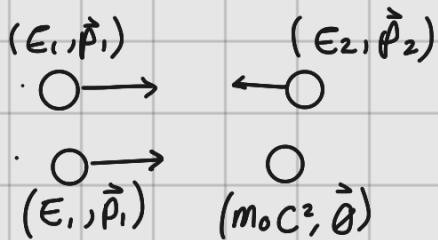
$BR_i$  = Branching ratio of a given decay

## COLLISIONS

- Rest mass is same in all frames

Frame:

- Centre of momentum:  $\vec{p}_1 = -\vec{p}_2$
- Fixed target:  $\vec{p}_2 = \vec{0}$
- Lab frame: Lab is at rest



- $S$  (Mandelstam- $S$ ) is invariant for collisions

$$S = (E_1 + E_2)^2 - |\vec{p}_1 + \vec{p}_2|^2 c^2 \rightarrow \text{for}$$

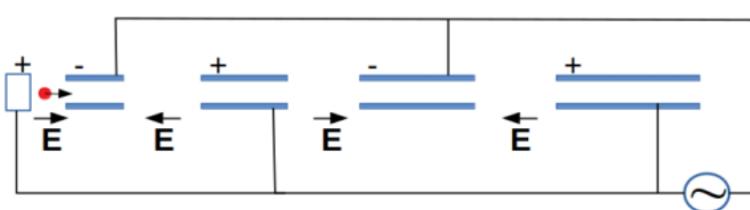
↳ In CM frame ( $\vec{p}_1 = -\vec{p}_2$ ):  $S_{\text{com}} = (E_1 + E_2)^2$

-  $\sqrt{S_{\text{com}}}$  = total energy available for particle creation

↳ In fixed target frame ( $\vec{p}_2 = \vec{0}$ )  $S_{\text{fixed}} = 2E_1 m_2 c^2 + m_1^2 c^4 + m_2^2 c^4$

# ACCELERATORS

## LINACS (linear accelerators)



particles accelerated through drift tubes of increasing length with an electric field between them

## Cyclotron

- Particle accelerated across gap between 2 electrodes
- Only for non-relativistic particles

$$r = \frac{mv}{qB}$$

## Synchrotron (eg. LHC)

- particles accelerated with fixed radius but increasing magnetic field
- Particles lose energy as photons (braking radiation)  $\Delta E \sim (\frac{E}{m})^4 \frac{1}{r}$   
 $\hookrightarrow r = \downarrow E \text{ loss} , \quad \uparrow m = \downarrow E \text{ loss}$

$$f_{\text{cyc}} = \frac{qB}{2\pi m}$$

$$f_{\text{cyc}} = \frac{qB}{2\pi m}$$

## DETECTORS

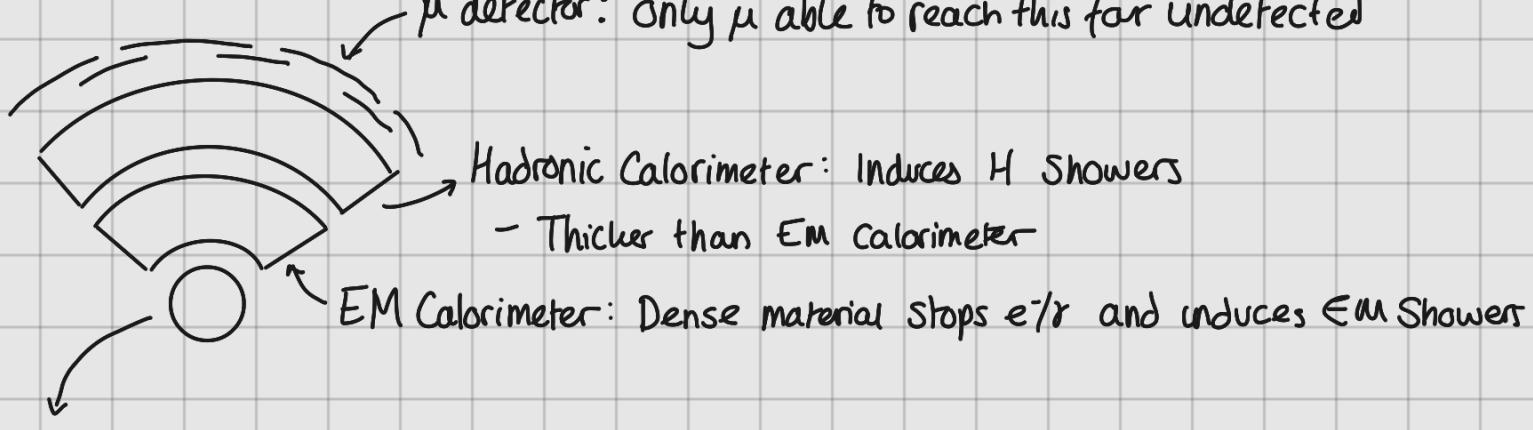
- $\mu$  deposit an  $E_{\text{min}}$  as they move through detector
- $e^-/\gamma$  initiate EM showers when undergoing EM interaction
- Hadrons initiate hadronic showers when involved in strong interaction

$x_0 = \text{radiation length}$

Probability of  $e^-/\gamma$  passing  $x$  through a material unaffected:  $P(x) = e^{-x/x_0}$

Probability of Hadron passing  $x$  through a material unaffected:  $P(x) = e^{-x/x_z}$

## Collider Detectors:



Tracker: Measures trajectory of colliding charged particles

- in a magnetic field

	Tracker	Ecal	HCal	$\mu$ chamber
$\mu$	→	→	→	→
$e^\pm$	→	⊗		
$\gamma$		⊗		
$H^\pm$	→	→	⊗	
$H^0$				⊗
$\nu$				

### PARTICLE SOURCES

- $\beta^-$  decay:  $n \rightarrow p + e^- + \bar{\nu}_e$
- Cosmic ray showers: Particles from astrophysical sources interact with atoms in atmosphere producing hadronic showers ( $\pi^0, k^0$  typically, then  $\pi^\pm \rightarrow \mu^\pm + \nu$ )
- Solar particles
  - Solar neutrinos from fusion in Sun
  - Solar wind: plasma of charged particles emitted from Solar Corona

### HELICITY

Helicity:

> helicity,  $h$  - the projection of Spin vector onto the direction of motion of a particle

$$h = \vec{S} \cdot \hat{p}$$

→  $\vec{S}$  doesn't change under  $\hat{P}$

- If  $h > 0$ , right-handed helicity
- If  $h < 0$ , left-handed helicity

→ Weak interaction only creates LH  $p$  and RH  $\bar{p}$

- Weak interaction violates  $\hat{C}\hat{P}$  — don't know why

→ there is a fundamental difference between physics of matter and antimatter

## DEFINITIONS

- > Spin - quantised intrinsic angular momentum
- > Weak isospin (unrelated above) - a quantum number relating to the electrically charged part of the weak interaction
- > Force - an exchange of quantum numbers by bosons
- > Principle of confinement - quarks can not be separated/isolated
- > Hadronization - separating 2 quarks until they create 2 more
- > Virtual particle - intermediate particles allowed to have any mass to make kinematics work
- > Symmetry - an operation on a system or the mathematics describing it that leaves the system or physics unchanged
- > Noether's theorem - Any system which displays continuous, differentiable symmetry embodies an associated conservation law
- > Axial vectors - vectors invariant to parity transforms

## 3 Principles of PP: Symmetry, Composition, Unification

### EQUATIONS

$$E^2 = (pc)^2 + (mc^2)^2 \quad E = hf \quad \lambda = \frac{h}{p} \quad h = \vec{s} \cdot \vec{p}$$

✓ helicity

$$N(t) = N(0) e^{-t/\tau}$$

$$S = (E_1 + E_2)^2 - |\vec{p}_1 + \vec{p}_2|^2 c^2$$

$$\Gamma = \frac{1}{\tau}$$

$$BR_i = \frac{\Gamma_i}{\Gamma_{\text{total}}}$$

$$S_{\text{com}} = (E_1 + E_2)^2$$

$$S_{\text{fixed}} = 2E_1 m_2 c^2 + m_1^2 c^4 + m_2^2 c^4$$

$$r = \frac{mv}{Bq}$$

$$f_{\text{cyc}} = \frac{qB}{2\pi rm}$$

$$\Delta E_{\text{or}} \sim \left(\frac{e}{m}\right)^4 \frac{1}{r}$$

$$P_{e/r}(x) = e^{-x/x_0}$$

$$P_H(x) = e^{-x/x_0}$$